# Spatial Extension of Quantum Gravitational Particles in $R^4 \times K^N$

## Gerald Rosen<sup>1</sup>

Received April 22, 1988

It is observed that the magnitude relation  $ma^2 \approx \hbar l_{\rm P}/c$  holds if the non-Euclidean incremental spatial volume associated with a fundamental particle of mass m and radius a is characteristically quantum gravitational in a Kaluza-Klein or superstring  $R^4 \times K^N$ . Here  $R^4$  is the four-dimensional Riemannian space-time of general relativity and  $K^N$  is a small-scale, compact, N-dimensional space of characteristic quantum gravitational volume  $l_{\rm P}^N$ , with  $l_{\rm p} = (\hbar G/c^3)^{1/2} = 1.61 \times 10^{-33}$  cm denoting the Planck length. For the electron and electron neutrino (assumed to possess nonzero mass bounded empirically by  $m_{\nu_e} \approx <30$  eV) the derived magnitude relation  $a \approx (\hbar l_{\rm P}/mc)^{1/2}$  yields the estimates  $a_e \approx 2.5 \times 10^{-22}$  cm and  $a_{\nu_e} \approx 3.3 \times 10^{-20}$  cm, spatial extensions which may be detectable by way of fine-scale effects in SSC experiments.

### **1. INTRODUCTION**

Recently it has been shown (Rosen, 1987) that the Einstein field equations

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = 8\pi G T_{\mu\nu} \tag{1}$$

are equivalent to the physicogeometric statement: A local spherical region of radius r containing energy with moment of inertia I has the physical volume

$$V(r) = (4\pi/3)(r^3 + GI)$$
(2)

in an associated free-falling (nonrotating, timelike geodesic) frame of reference. Here Newton's constant is  $G = 7.41 \times 10^{-29}$  cm/g in a system of physical units such that the speed of light in vacuum equals unity (c = 1). In particular, for a uniform spherical distribution of radius a and total energy

<sup>1</sup>Department of Physics, Drexel University, Philadelphia, Pennsylvania 19104.

1351

m, one has  $I = \frac{2}{5}ma^2$ ; therefore, the non-Euclidean incremental spatial volume  $\Delta V \equiv V(a) - 4\pi a^3/3$ , i.e.,

$$\Delta V = (8\pi/15) Gma^2 \tag{3}$$

is attributed to a mass m of radius a, according to general relativity and formula (2). Conversely, if (2) or (3) is valid for any spherical region of arbitrary radius containing a uniform distribution of energy in a free-falling frame of reference, then the gravitational field equations (1) must hold through the space-time region. The non-Euclidean character of physical space, as prescribed by general relativity, is expressed in an equivalent quantitative fashion by (3).

## 2. QUANTUM GRAVITATIONAL PARTICLES IN $R^4 \times K^N$

In contemporary studies of the Kaluza-Klein and superstring models (Lovelace, 1984; Fradkin and Tseytlin, 1985; Ne'eman and Sijacki, 1986), space-time is viewed as a (4+N)-dimensional manifold with the structure  $R^4 \times K^N$ , the direct product of the four-dimensional Riemannian space-time of general relativity and a small-scale, compact, N-dimensional space of characteristic quantum gravitational volume  $l_{\rm P}^N$ , where  $l_{\rm P} \equiv (\hbar G)^{1/2} =$  $1.61 \times 10^{-33}$  cm is the Planck length. The physical distinction between the three spatial dimensions of  $R^4$  and the N spatial dimensions of  $K^N$  is likely to become indefinite on a scale of order  $l_{\rm P}$ , due to quantum gravitational fluctuations or a possible discrete aspect to space-time that underlies spatial dimensionality [see, for example, Bombelli et al. (1987) and work cited therein]. If all N+3 spatial dimensions enter on the same footing at the smallest physical scales, then the small-scale equivalence of the N+3spatial dimensions suggests that the non-Euclidean incremental (N +3)-dimensional spatial volume attributed to a fundamental particle of mass m and radius a is also characteristically quantum gravitational in (a t = consthypersurface of)  $R^4 \times K^N$ . If this quantum gravitational feature is indeed manifest at the smallest spatial scales, it follows that the (N+3)-dimensional volume must be of order  $l_{\rm P}^{N+3}$  and the magnitude relation  $\Delta V \sim l_{\rm P}^3$  must hold for the quantity (3) in  $R^4$ . Thus, such a characteristically quantum gravitational particle in  $R^4 \times K^N$  requires the magnitude relation  $Gma^2 \approx l_P^3$ . or equivalently

$$ma^2 \sim \hbar l_{\rm P}$$
 (4)

to hold as a quantum constraint on the product of the particle mass times its radius squared.

For the electron and electron neutrino (assumed to possess nonzero mass bounded empirically by  $m_{\nu_e} \approx 30 \text{ eV}$ ) the spatial extension estimate

#### **Quantum Gravitational Particles**

that follows from (4),

$$a \approx (\hbar l_{\rm P}/m)^{1/2} \tag{5}$$

yields

$$a_e \sim 2.5 \times 10^{-22} \,\mathrm{cm}$$
 (6)

$$a_{\nu_a} \approx 3.3 \times 10^{-20} \,\mathrm{cm}$$
 (7)

The latter very small spatial extensions may be detectable by way of finite-scale effects in SSC experiments, with the lengths on the right side of (6) and (7) giving the scales at which measurable departures from the predictions of standard (point-charge) QED and electroweak theory may be expected to show up.

### REFERENCES

Bombelli, L., et al. (1987). Physical Review Letters, 59, 521-525.
Fradkin, W. S., and Tseytlin, A. A. (1985). Physical Letters, 158B, 316-321.
Lovelace, C. (1984). Physics Letters, 135B, 75-79.
Ne'eman, Y., and Sijacki, D. (1986). Physics Letters, 174B, 165-169.
Rosen, G. (1987). Physical Review D, 35, 2595-2598.